

F. The Fermi Temperature T_F sets Fermi Gas' Temperature Scale

- $T = 0K$ Physics $\Rightarrow$ fermions stack up in s.p. states $\Rightarrow$ filled up to $\epsilon = E_F$

$$E_F = \frac{\hbar^2}{2m} \left(3\pi^2 \frac{N}{V} \right)^{2/3} \equiv kT_F \Rightarrow T_F = \frac{1}{k} \frac{\hbar^2}{2m} \left(3\pi^2 \frac{N}{V} \right)^{2/3} \sim \left(\frac{N}{V} \right)^{2/3}$$

[Metals: $E_F \sim \text{few eV}$, $T_F \sim \text{few } 10^4 K$]

so metal physics is low-temperature physics

Key Physics is that Fermions observe the Pauli Exclusion Rule

or wavefunction requirement in QM

- When would fermions don't need to care about their Quantum character?

We know that all gases (fermi/Bose gases) approaches classical ideal gas at high temperature and dilute (small $\frac{N}{V}$) limits

Recall: Classical Ideal Gas behavior requires (see Ch. IX)

$$\left(\frac{V}{N}\right) \frac{1}{\lambda_{th}^3(T)} \gg 1 \quad \text{OR} \quad \underbrace{\left(\frac{V}{N}\right)^{1/3}}_{\text{spacing between particles in system}} \gg \underbrace{\lambda_{th}(T)}_{\substack{h \\ \sqrt{2\pi mkT} \\ \text{thermal} \\ \text{de Broglie wavelength}}} \quad (20)$$

Particles NOT crowded

⇒ No need to worry about quantum nature

The criterion also sets a criterion on temperature

$$\frac{h}{\sqrt{2\pi mkT_0}} \approx \left(\frac{V}{N}\right)^{1/3}$$

$$h \left(\frac{N}{V}\right)^{1/3} \approx \sqrt{2\pi mkT_0}$$

$$\text{OR } T_0 \sim \frac{h^2}{2\pi m k} \frac{1}{\left(\frac{V}{N}\right)^{2/3}} \quad (21)$$

we need to worry about quantum nature of particles when $T \lesssim T_0$

For Fermions, $T_0 \approx T_F$

for electrons (m_e is tiny), $T_F \sim \frac{\hbar^2}{2m_e} \left(3\pi^2 \frac{N}{V} \right)^{2/3}$ very high temperature
 ($\frac{N}{V}$ is high in metals)

(i) T_F is also the temperature below which quantum nature of fermions matters (this is why we need to use $f_{FD}(\epsilon)$ to study Fermi Gas physics)

(ii) Only when $T > T_F$ (or $T \gg T_F$), behavior of Fermi Gas will approach⁺ that of classical gas

[note: metals would have melted at such temperatures]

⁺ See next section for behavior of Fermi Gas for $T \gg T_F$

For Bosons

- Argument in getting $T_0 \approx \frac{h^2}{2m\pi k} \left(\frac{N}{V}\right)^{2/3}$ does NOT rely on fermions or bosons

- Some atoms are bosons

nucleus + electrons $\left\{ \begin{array}{l} m \rightarrow m_{\text{atom}} \gg m_e \text{ (} 10^3 \text{ times higher)} \\ \text{atom gases are dilute } \left(\frac{N}{V}\right) \ll n_{\text{electron in metals}} \end{array} \right.$

$\therefore T_0^{(\text{bosons})}$ is typically very low

$T < T_0^{(\text{bosons})} \Rightarrow$ Bosonic nature comes in
(Bose-Einstein Condensation!)

$\hookrightarrow$ hard to observe because $T_0^{(\text{boson})}$ is low
and most gases would have turned into liquids/solids

G. Ideal Fermi Gas ain't classical ideal gas even at high temperature

[This section is Optional]

Question: What is the behavior of Ideal Fermi Gas at high temperature?

What for? All gases approach classical ideal gas at high temperatures, but any signature of fermions in high-T behavior?

Classical ideal gas $pV = NkT$ or $\frac{p}{kT} = \frac{N}{V}$

Fermi Gas at high temperature $\frac{pV}{NkT} = 1 + (\text{something})?$

or $\frac{p}{kT} = \frac{N}{V} + \underbrace{(\text{some factor})}_{\text{What is this?}} \left(\frac{N}{V}\right)^2$

Approach

▪ We have Eqs. (1), (2), (3), (4) [Sec. A] $g_s = 2$ used

▪ Eq. (4) says $pV = \frac{2}{3} E = \frac{2}{3} \underbrace{\frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2}}_A \int_0^\infty \frac{\epsilon^{3/2}}{e^{(\epsilon-\mu)/kT} + 1} d\epsilon$

[Call $x = \epsilon/kT$, $\epsilon^{3/2} = (kT)^{3/2} x^{3/2}$, $d\epsilon = (kT) dx$, $\epsilon^{1/2} = (kT)^{1/2} x^{1/2}$]

$$pV = \frac{2}{3} A (kT)^{5/2} \underbrace{\int_0^\infty \frac{x^{3/2}}{e^{-\mu/kT} e^x + 1} dx}_{\text{after integrating over } x, \text{ it is a function of } e^{-\mu/kT}} \quad (22) \quad (\text{Exact})$$

Similarly, Eq. (1) says

$$N = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} \int_0^\infty \frac{\epsilon^{1/2}}{e^{(\epsilon-\mu)/kT} + 1} d\epsilon = A (kT)^{3/2} \underbrace{\int_0^\infty \frac{x^{1/2}}{e^{-\mu/kT} e^x + 1} dx}_{\text{after integrating over } x, \text{ it is a function of } e^{-\mu/kT}} \quad (23) \quad (\text{Exact})$$

after integrating over x , it is a function of $e^{-\mu/kT}$

Let's call $e^{+\mu/kT} \equiv \xi$, so $e^{-\mu/kT} = \xi^{-1}$ (24)

Define $\int_0^\infty \frac{x^{n-1}}{\xi^{-1} e^x + 1} dx = \underbrace{\Gamma(n)}_{\text{Gamma Function}^+} \cdot f_n(\xi)$ (25)

Using Eqs. (22), (23),

$$\frac{pV}{N} = \frac{2}{3} kT \frac{\int_0^\infty \frac{x^{3/2}}{\xi^{-1} e^x + 1} dx}{\int_0^\infty \frac{x^{1/2}}{\xi^{-1} e^x + 1} dx}$$

$$\frac{pV}{NkT} = \frac{2}{3} \frac{\Gamma(5/2) f_{5/2}(\xi)}{\Gamma(3/2) f_{3/2}(\xi)} = \frac{2}{3} \frac{\cancel{\frac{3}{4}\sqrt{\pi}}}{\frac{1}{2}\sqrt{\pi}} \frac{f_{5/2}(\xi)}{f_{3/2}(\xi)} = \frac{f_{5/2}(\xi)}{f_{3/2}(\xi)}$$

this is "1"

$$\Rightarrow \boxed{\frac{pV}{NkT} = \frac{f_{5/2}(\xi)}{f_{3/2}(\xi)}} \quad (26)$$

Exact (True for all temperatures)

Note⁺:

$$\Gamma(1/2) = \sqrt{\pi}$$

$$\text{as } \Gamma(n) = \int_0^\infty x^{n-1} e^{-x} dx$$

$$\Gamma(3/2) = \frac{1}{2} \Gamma(1/2) = \frac{\sqrt{\pi}}{2}$$

$$\Gamma(5/2) = \frac{3}{2} \Gamma(3/2) = \frac{3}{4} \sqrt{\pi}$$

⁺ See Notes on the Gamma Functions under "Essential Math Skills"

So far, all equations are exact, just rewriting Eqs. (1) - (4).

If we know ξ ($\xi \equiv e^{\mu/kT}$ (but μ itself can be negative as in classical gas)),
then Eq. (26) gives $\frac{pV}{NkT}$ (more than just "1").

Recall $\mu(T)$ (thus $\xi = e^{\mu/kT}$) is fixed by the "N-equation" (Eq. (1) & Eq. (23))

i.e.

$$N = \frac{V}{2\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} (kT)^{3/2} \underbrace{I^{3/2}}_{\frac{\sqrt{\pi}}{2}} f_{3/2}(\xi)$$

$$= \underbrace{g_s}_{\text{spin degeneracy}} \frac{V}{4\pi^2} \left(\frac{\sqrt{2mkT}}{\hbar} \right)^3 \cdot \frac{\sqrt{\pi}}{2} f_{3/2}(\xi)$$

$$= 2 V \left(\frac{\sqrt{2\pi mkT}}{\hbar} \right)^3 f_{3/2}(\xi) = 2 \frac{V}{\lambda_{th}^3(T)} f_{3/2}(\xi)$$

$$\boxed{N = 2 \frac{V}{\lambda_{th}^3} f_{3/2}(\xi)} \quad (27) \text{ (Exact) can be used to fix } \xi$$

Summary-

$$\frac{pV}{NkT} = \frac{f_{5/2}(\xi)}{f_{3/2}(\xi)} ; N = 2 \frac{V}{\lambda_{th}^3} f_{3/2}(\xi) ; \xi = e^{\mu/kT}$$

advanced-level way of writing the governing equations of 3D ideal Fermi gas

$$f_n(\xi) = \frac{1}{\Gamma(n)} \int_0^\infty \frac{x^{n-1}}{\xi^{-1} e^x - 1} dx$$

They are exact, and applicable to all temperature.

Back to our goal, what is behavior in approach the classical gas limit?

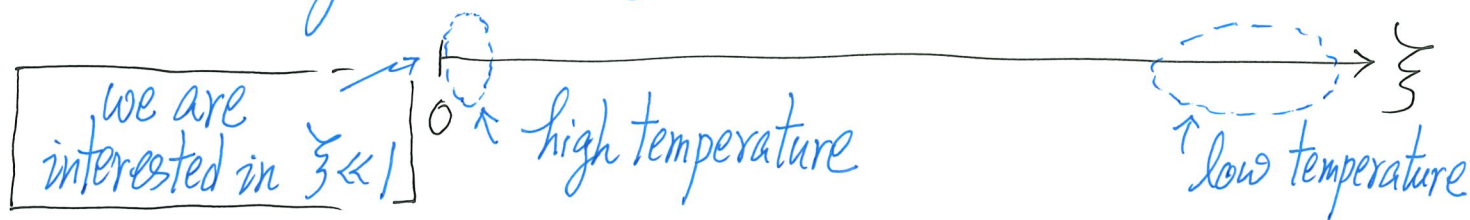
Recall: Obtained for classical ideal gas $\mu = -kT \ln \left[\frac{V}{N} \left(\frac{\sqrt{2\pi m kT}}{h} \right)^3 \right]$
 [see Application CE-②]

$$\therefore \zeta_{\text{classical}} = e^{\mu/kT} = e^{-\ln \left[\frac{V}{N} \frac{1}{\lambda_{th}^3} \right]} = \frac{\lambda_{th}^3}{\left(\frac{V}{N} \right)} \ll 1$$

negative $\nearrow$ large $\left(\frac{V}{N} \right)^{1/3} \gg \lambda_{th}^3$ (ideal gas)

As we want to study ideal Fermi Gas as it approaches the classical gas limit, we expect $\zeta \approx \zeta_{\text{classical}} \ll 1$ (28) ($T \gg T_F$ situation)

Formally, $\mu(T=0) = E_F$, $\zeta_{(T=0)} = e^{E_F/kT} \rightarrow \infty$; low- T , $\mu(T)$ shifts down a bit, ζ is large.
 i.e. for ideal Fermi Gas $0 < \zeta < \infty$



Strategy Update : Study $f_{5/2}(\xi)$ and $f_{3/2}(\xi)$ for $\xi \ll 1$

small parameter for the regime we are interested in

$$\int_0^\infty \frac{x^{n-1}}{\xi^{-1} e^x + 1} dx = \int_0^\infty \frac{\xi x^{n-1} e^{-x}}{1 + \xi e^{-x}} dx$$

(Exact)

$$\approx \int_0^\infty \xi x^{n-1} e^{-x} \left[1 - \xi e^{-x} + \underbrace{(\xi^2 \text{ and higher terms})}_{\text{ignored } (\xi \ll 1)} \right] dx$$

$$= \xi \int_0^\infty x^{n-1} e^{-x} dx - \xi^2 \int_0^\infty x^{n-1} e^{-2x} dx$$

($y=2x$ in 2nd term)

$$= \xi I^1(n) - \xi^2 \int_0^\infty \frac{y^{n-1}}{2^{n-1}} e^{-y} \frac{dy}{2}$$

for $\xi \ll 1$

$$= \xi I^1(n) - \frac{\xi^2}{2^n} I^1(n) = I^1(n) \left[\xi - \frac{\xi^2}{2^n} \right] = I^1(n) \overbrace{\left[\xi - \frac{\xi^2}{2^n} \right]}^{f_n(\xi)}$$

this is $f_n(\xi)$ for $\xi \ll 1$

$$\therefore \boxed{f_n(\xi) = \xi - \frac{\xi^2}{2^n}} \quad (29) \text{ for } \xi \ll 1$$

Recall: μ (or ξ now) is fixed by Eq. (1).

$$N = \underbrace{(2)}_{\text{spin degeneracy (usually not included in classical ideal gas calculations)}} \frac{V}{\lambda_{th}^3(T)} f_{3/2}(\xi) \approx 2 \frac{V}{\lambda_{th}^3} \left(\xi - \frac{\xi^2}{2^{3/2}} \right)$$

$$\Rightarrow \xi \approx \underbrace{\left(\frac{N}{V} \frac{\lambda_{th}^3}{2} \right)}_{\text{this is } \left(\frac{1}{2} \xi_{\text{classical}} \right) < 1, \text{ this is the leading term}} + \frac{\xi^2}{2^{3/2}}$$

$$\approx \left(\frac{N}{V} \frac{\lambda_{th}^3}{2} \right) + \underbrace{\frac{1}{2^{3/2}} \left(\frac{N}{V} \frac{\lambda_{th}^3}{2} \right)^2}_{\text{Even smaller!}}$$

$$\approx \underbrace{\left(\frac{N}{V} \frac{\lambda_{th}^3}{2} \right)}_{\text{just the classical ideal gas value}} \quad (30)$$

sufficient for our purpose

$$(\text{temperature } T \text{ inside } \lambda_{th} = \frac{h}{\sqrt{2\pi m k T}})$$

$$\text{Eq. (26): } \frac{pV}{NkT} = \frac{f_{5/2}(\xi)}{f_{3/2}(\xi)} \quad (\text{exact}) \approx \frac{\xi - \frac{\xi^2}{2^{5/2}}}{\xi - \frac{\xi^2}{2^{3/2}}} \quad (\xi \ll 1, \text{ approaching classical gas limit})$$

$$\approx \left(1 - \frac{\xi}{2^{5/2}}\right) \left(1 + \frac{\xi}{2^{3/2}}\right) \approx 1 + \xi \left(\frac{1}{2^{3/2}} - \frac{1}{2^{5/2}}\right) + \text{ignored } (\xi^2, \xi^3 \dots)$$

$$= 1 + \frac{1}{4\sqrt{2}} \xi$$

$$= 1 + \frac{1}{4\sqrt{2}} \frac{\lambda_{th}^3}{2} \left(\frac{N}{V}\right)$$

using ξ in Eq. (30)
(31)

$$\therefore \boxed{\frac{p}{kT} = \frac{N}{V} + \frac{1}{4\sqrt{2}} \frac{\lambda_{th}^3}{2} \left(\frac{N}{V}\right)^2 = \frac{N}{V} + B_2(T) \left(\frac{N}{V}\right)^2} \quad (32)$$

something!

$\therefore$ Ideal (Non-interacting) Fermi Gas behaves as $\frac{p}{kT} = n + B_2(T)n^2$ with a Positive Second Virial Coefficient $B_2(T) = \frac{1}{4\sqrt{2}} \frac{\lambda_{th}^3}{2}$

Positive $B_2(T)$ implies some effective repulsion between fermions, even though they don't interact!

Aside : We saw in Van der Waals Gas $(P + \frac{n^2 a}{V^2})(V - nb) = nRT$

attractive
 ↗
 repulsive
 ↗
 # moles

that

$$\frac{p}{kT} = n + \underbrace{B_2(T)}_{\left[+ \frac{b}{N_A} - \frac{a}{N_A^2 kT} \right]} n^2$$

attractive part contributes negative term to B_2
 ↗
 repulsive part of particle-particle
 interaction contributes positive term to B_2

Summary

- Even when ideal Fermi Gas approaches the classical ideal gas limit, the first correction term indicates the quantum nature of the fermions with a positive B_2 , as if they have an effective repulsive interaction.

Cross Reference : There is an analogous treatment for Ideal Bose gas with $B_2(T) = -\frac{1}{4\sqrt{2}} \frac{\lambda_{th}^3}{g_s}$, i.e. negative $B_2(T)$.

References

- Fermi Gas is a fascinating subject covered in most statistical mechanics books
- See Pathria, "Statistical Mechanics", Ch. 8

Greiner, Neise, Stöcker, "Thermodynamics and Statistical Mechanics", Ch. 14

[both provided a discussion on white dwarf stars]

- Ashcroft and Mermin, "Solid State Physics", Ch. 1, 2, 3

[Sommerfeld model of metals and its failure]

Beyond Ideal Fermi Gas : Interacting Systems

Pathria (see above), Ch. 11

Diep, "Statistical Physics", Ch. 7 (Method of Second Quantization),

Ch. 10 (Systems of Interacting Electrons: Fermi Liquids)